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Quasi-geostrophic turbulence and generalized scale invariance, a theoretical reply

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Abstract. Lindborg et al. (2010) claim that the apparent spectrum power law $E(k) \approx k^{-3}$ on scales $\geq 600$ km obtained with the help of commercial jetliner trajectory deviations (GASP and Mozaic databases) could not be brought into question (Lovejoy et al., 2009a), because this spectrum corresponds to “a well known theory of quasi-geostrophic turbulence developed by Charney (1971)”. Lindborg et al. (2010) also claim that “limitations [of this theory] have been relaxed in many of the modern models of atmospheric turbulence”. We show that both claims are irrelevant and that generalized scale invariance (GSI) is indispensable to go beyond the quasi-geostrophic limitations, to go in fact from scale analysis to scaling analysis in order to derive better analytical models. In this direction, we derive vorticity equations in a space of (fractal) dimension $D = 2 + H_z$ ($0 \leq H_z \leq 1$), which corresponds to a first step in the derivation of a dynamical alternative to the quasi-geostrophic approximation and turbulence. The corresponding precise definition of fractional dimensional turbulence already demonstrates that the classical 2-D and 3-D turbulence are not the main options to understand atmospheric dynamics. Although $(2 + H_z)$-D turbulence (with $0 < H_z < 1$) has more common features with 3-D turbulence than with 2-D turbulence, it has nevertheless very distinctive features: its scaling anisotropy is in agreement with the layered pancake structure, which is typical of rotating and stratified turbulence but not of the classical 3-D turbulence.

1 Introduction

The scale analysis methodology developed by Charney (1948) in his seminal derivation of the quasi-geostrophic (QG) approximation has become standard in meteorology and oceanography (Pedlosky, 1979). Unfortunately, its common and distinguishing features with respect to a scaling analysis have not yet been fully recognized. It is ironical that the source of the present debate on intermediate scale atmospheric dynamics presumably corresponds to the importation into meteorology of two successive techniques from hydrodynamics at two different periods. In order to focus the present paper on this question and therefore on the limitations of the theory of quasi-geostrophic turbulence (QGT, Charney, 1971) and show how to overcome them, let us first reject the second claim of Lindborg et al. (2010), LTNGC hereafter, that models may overcome limitations of a theory because in a very general manner models are obtained by introducing further constraints into a given theoretical framework, e.g. boundary conditions, discretization of partial differential equations, subgrid modelling and other parametrizations. Furthermore, these constraints in the unique quasi-geostrophic model simulation (Tung and Orlando, 2003) cited by LTNGC were such a problem that, as discussed below, they seem to have introduced spurious numerical estimates of the scaling ranges, instead of relaxing limitations of the theory.

Before addressing QGT, we will first discuss a few fundamental features of the derivation of the QG approximation itself, as well as its motivation. This detailed discussion (Sect. 2) is necessary (i) to better evaluate the limitations of quasi-geostrophic turbulence (QGT, Sect. 3) and (ii) to enable us to derive an alternative (Sect. 4). In particular,
we show that to better understand the fundamentals of atmospheric dynamics a scaling analysis is required instead of a scale analysis, which was the basic technique used by Charney in his derivation and led to a drastic reduction of vortex stretching. Both questions are discussed throughout our paper and they enable us to obtain in a straightforward manner a vorticity equation for a turbulence that at first glance looks two-dimensional at large scales, three-dimensional at small scales, but is in fact of (fractal) dimension \( D = 2 + H_z \) \( (0 \leq H_z \leq 1) \) at all scales as illustrated by Fig. 1. We believe that this equation may satisfy those who asked for a dynamical alternative to QGT (Yano, 2010), although this should be completed by a similar treatment of the thermodynamic energy evolution equation.

As our paper is already long, we prefer to acknowledge the necessity of this second step rather than to proceed with it. It may be helpful to note that this paper does correspond to a reply to the comment issued by LTNGC on the paper of Lovejoy et al. (2009a), LTSH hereafter. However, because the debate with LTNCG bears on very fundamental issues and not on technical details this yielded a stand alone paper that has many more original results (e.g. new dynamical equations, as well an original technique to derive them) than a usual reply to a comment. As mentioned in the title, it is a “theoretical reply”, because it is focused on theory, not on the empirical evidence brought by LTSH that the empirical spectral slope value is closer to 2.2 (as for buoyancy subrange) than to 3 (as for the enstrophy (i.e. square vorticity) inertial subrange of 2-D turbulence) and could correspond to the fact that aircraft sample the vertical fluctuations instead of the horizontal ones. The present paper is indeed focused on the dynamical/physical modelling of these observations.

2 QG vorticity equation and scale analysis

Before discussing the derivation of the QG approximation, let us highlight why this approximation has remained so attractive and therefore popular. Historically, it was the first mathematically self-consistent derivation of a closed dynamical system from the diagnostic geostrophic and hydrostatic approximations. It also respects the conservation laws of potential temperature and absolute potential vorticity, at least their corresponding first order expressions within the QG approximation. Furthermore, the global mathematical properties (e.g. for large times) of QG solutions are known, unlike those of the Navier-Stokes equations, as well as the fact that exact results were obtained on their order of approximation with respect to the primitive equations (PE) with rather mild restrictions on the initial conditions (Bourgeois and Beale, 1994). However, PE correspond already to a given set of approximations (e.g. inviscid, incompressible flow with variable density, Boussinesq and (quasi-) hydrostatic assumptions, (Pedlosky, 1979) for discussion), which can be brought into question on various ranges of scale.

Let us also recall that the derivation by Charney (1948) of the QG approximation was based on a scale analysis, borrowed from aerodynamic boundary layer theory (Charney refers to Goldstein, 1938) to intentionally filter out the “meteorologically insignificant wave components” from the “meteorologically significant motions”. The latter were considered to be distinguishable from all other types of atmospheric motion only by a great difference in scales. This question of filtering was essential for the sake of the pioneering development of numerical weather forecasting (Charney et al., 1950) and a part of the paper by Charney (1948) was indeed devoted to checking that this filtering was efficient for acoustic waves. Nevertheless, the more general question of which structures were filtered out was not addressed. Another important requirement put forward by Charney was also related to numerical forecasts: the approximation should be (easily) computable. It is worthwhile noting that none of these issues is related to the question of obtaining a rather complete description of the atmospheric dynamics.

More precisely, the scale analysis performed by Charney was based on the common idea in meteorological practice that large scale dynamics were quasi-two-dimensional, quasi-geostrophic and quasi-hydrostatic, due to the fact that the rotation \( \Omega \) and the gravitational field \( g \) of the Earth introduce preferential directions in the Navier-Stokes equations:

\[
\frac{Du}{Dt} + 2\Omega \times u = -\nabla p/\rho + g + v \Delta u
\]

where \( u \) is the velocity, \( p \) the pressure, \( \rho \) the density, \( v \) the viscosity, \( \frac{Du}{Dt} \) denotes the material derivative, i.e. the time derivative following the motion \( (\partial/\partial t + u \cdot \nabla) \), \( \nabla \) denotes the 3-D gradient operator, \( \Delta = \nabla^2 \) the 3-D Laplace operator. The velocity field is assumed to be solenoidal (div\( u \) = 0), so the mass conservation equation reads:

\[
\frac{D\rho}{Dt} = 0
\]
Although the assumption of incompressibility introduces limitations, it can be argued that this assumption is justifiable for small scales, and pressure coordinates may avoid this issue for large scales. There is no fundamental difficulty in dealing with these equations in spherical coordinates, or more generally in local manifold coordinates, however it is usual to approximate Eq. (1) on the tangent space to the Earth at a given point of longitude \( \phi_0 \) and latitude \( \theta_0 \), with a vertical unit vector \( \mathbf{n} \):

\[
Du/ Dt + f \mathbf{n} \times \mathbf{u} = -\nabla p/\rho + g + \nu \Delta \mathbf{u} \tag{3}
\]

where only the vertical component \( f \mathbf{n} \) of the Earth’s angular velocity is taken into account, \( f \) being the Coriolis parameter at the latitude \( \theta_0 \). Furthermore, the variation of \( f \) with respect to the south-north coordinate \( \gamma = a(\theta - \theta_0) \) (\( a \) being the Earth radius), is usually linearized (the so-called \( \beta \) ‐plane approximation):

\[
f(\gamma) = f_0 + \beta \gamma; \; (f_0, \beta) = 2\Omega (\sin \theta_0, \cos \theta_0) \tag{4}
\]

The scales of interest being much larger than those of the dissipation range, the corresponding term will be omitted in the following (without excluding an effective dissipation at small scales), i.e. the viscosity is considered as infinitely small or the Reynolds number \( Re = UL/\nu = ZL^2/\nu \) and the Ekman number \( E = gL/\Omega \) as infinitely large, where \( Z \) denotes the (relative) vorticity scale of the atmosphere, \( L \) the outer scale of the horizontal fluctuations, \( U \) the corresponding scale of the (horizontal) velocity. For infinitely small Rossby numbers \( Ro = Z/\Omega \), the Navier-Stokes equations reduce to the (diagnostic) geostrophic balance equation with the corresponding 2-D geostrophic solution:

\[
2\Omega \times \mathbf{u}_g = -\nabla p/\rho + g \Rightarrow \mathbf{u}_g = \Omega \times (\mathbf{g} - \nabla (p/\rho))/(2\Omega^2) \tag{5}
\]

The latter can furthermore be approximated with the help of Eq. (3) and of the stream function \( \psi \) as:

\[
\mathbf{u}_g = \mathbf{n} \times \nabla \psi = \nabla \times \mathbf{n} \psi; \; \psi = (p/\rho_0 - gz)/f_0 \tag{6}
\]

the sub-index \( 0 \) of the (barotropic) hydrostatic density \( \rho_0 \), as for other variables, refers to a reference or background vertical profile, which can be understood as a time average of instantaneous profiles. The vertical variation of the stream function corresponds to:

\[
\partial \psi/\partial z = \partial (p - \rho_0)/\rho_0 f_0 \partial z = -\partial \rho/\rho_0 f_0 \partial z \tag{7}
\]

The \( \mathrm{QG} \) approximation corresponds to introducing a linear three-dimensional ageostrophic perturbation \( \mathbf{u}_{ag, w} \) of order \( Ro \) to the zero-order two-dimensional geostrophic solution \( \mathbf{u}_g, 0 \) and similarly for the relative vorticity \( \xi \):

\[
\mathbf{u} = (\mathbf{u}_h, w) \approx (\mathbf{u}_g, 0) + (\mathbf{u}_{ag, w}); \; \xi = (0, \zeta_g) + \zeta_{ag} \tag{8}
\]

The material derivative is systematically approximated by its geostrophic approximation (\( \nabla_h \) denotes the horizontal gradient):

\[
D_g/ Dt = \partial/ \partial t + \mathbf{u}_g \cdot \nabla_h \tag{9}
\]

which, as emphasized by Charney, is much more manageable than the horizontal material derivative \( (D_h/ Dt = \partial/\partial t + \mathbf{u}_h \cdot \nabla_h) \). In particular, \( D_g/ Dt \) is easily defined with the help of the stream function \( \psi \), (\( J \) denotes the 2-D horizontal Jacobian):

\[
D_g/ Dt = \partial/\partial t + J(\psi,..) \tag{10}
\]

To derive the \( \mathrm{QG} \) approximation, it is convenient to start from the vorticity equation (the curl of Eq. (1), \( \omega = 2\Omega + \xi \) denotes the total vorticity):

\[
D\omega/ Dt = \mathbf{s} + \mathbf{b}; \; \mathbf{s} = (\omega \cdot \nabla)\mathbf{u}; \; \mathbf{b} = \nabla \rho \times \nabla (p/\rho)^2 \tag{11}
\]

We respectively call the vectors \( \mathbf{s} \) and \( \mathbf{b} \) stretching and baroclinic vectors, because the main action of the former is to stretch the vortex field, although it also has the important role of tilting the vorticity from one direction to another, and the latter is both the source and a measure of the flow baroclinicity (e.g. it is obviously zero for barotropic flows, where \( \nabla \rho \) and \( \nabla p \) are collinear). Whereas the baroclinic vector \( \mathbf{b} \) is of second order in a quasi-barotropic flow, the stretching vector \( \mathbf{s} \) is of first order and non-zero contrary to strictly 2-D fluid motions (including geostrophic motions), because the (vertical) vorticity and the (horizontal) velocity gradient would be orthogonal in this case. For \( \mathrm{QG} \), the leading term is obtained as the product of the (small) vertical gradient of the velocity with the (large vertical) vorticity component:

\[
(\zeta + f \mathbf{n}) \cdot \nabla u \approx f \partial w/\partial z \approx f_0 \partial w/\partial z \tag{12}
\]

We consider that usually this approximation is not sufficiently discussed. Whereas it has the interest to preserve a non-zero term, as in 3-D turbulence and contrary to 2-D turbulence, the linear assumption inherent in the \( \mathrm{QG} \) stretching term (Eq. (12) unfortunately suppresses nonlinear growth of vorticity, which is the mechanism that can generate a catastrophe/divergence of enstrophy in a finite time and gives the possibility of a non-zero energy dissipation for an infinitesimally small viscosity in 3-D turbulence (e.g. Lesieur, 2008). The resulting \( \mathrm{QG} \) vorticity equation corresponds to:

\[
D_g(\zeta_g + \beta \gamma)/ Dt = f_0 \partial w/\partial z \tag{13}
\]

Together with a linearisation of the thermodynamic energy evolution equation (briefly discussed below), this yields the conservation law of the quantity \( q \) called the pseudo potential vorticity (or \( \mathrm{QG} \) potential vorticity) by analogy to Ertel’s theorem (Ertel, 1942, for review: Mueller, 1995):

\[
D_g q/ Dt = 0; \; q = \zeta + \beta \gamma - g f_0 \partial (\rho/\rho_0^2)/(\partial \zeta_g + \zeta_{ag}) \tag{14}
\]

As explained in the next section, this conservation law is fundamental for \( \mathrm{QG} \) dynamics and its scaling. This law is obtained by eliminating the vertical velocity \( w \) from the vorticity equation (Eq. (13) with the help of the mass conservation linearly approximated by:

\[
D_g \rho/ Dt - w N^2 \rho_0/ g = 0; \; N^2 = gd \log (\theta_0)/dz \tag{15}
\]
where $N$ is the (mean) Brunt-Väisälä frequency defined with the help of the vertical profile of the background potential temperature $\theta_0$, which is assumed to be smooth, i.e. no important temperature gradients along the horizontal. It is remarkable that Charney (1971) – as emphasized by Schertzer and Lovejoy (1985a) – readily admitted the limitations of the QG approximation and expressed the question of separation of scales in terms of temperature gradients that must remain quite moderate.

One reason for the enormous success of the QG approximation is that the pseudo potential vorticity (Eq. 14) can be rewritten with only the help of the stream function $\psi$ (due to Eq. 7):

$$ q = \Delta \psi + \beta y + \partial(f_0^2 \partial \psi / N^2 \partial z) / \partial z $$

which together with Eqs. (10), (14) yields a compact dynamical equation for the stream function:

$$ (\partial/\partial t + J(\psi, .))(\Delta \psi + \beta y + \partial(f_0^2 \partial \psi / N^2 \partial z) / \partial z) = 0 $$

3 Scaling analysis of the QG approximation and QG turbulence

The domain of validity of the QG approximation is a priori restricted to intermediate scales (relatively small Rossby numbers), because a series of assumptions bring into question the QG validity on large scales (infinitely small Rossby numbers: quasi-uniform reference state density, simplified representation of the Coriolis force, Eq. 4) whereas other assumptions (linearization of the stretching vector and of the thermodynamic energy evolution equation) do it at small scales (large Rossby numbers). However, this does not prevent the possibility of studying the scaling behaviour of its solutions over a wider scale range. In particular, one can always argue (Yano, 2011) that the a posteriori domain of validity of a given approximation can be wider than the validity domain of its original derivation. However, this cannot be overstated, because there is in the present case not only the lack of empirical evidence (as discussed by Lovejoy et al., 2009a), but also the general fact that a linear approximation of a nonlinear system like QG may hold only over a limited range of scale. In the present case, this requires that small scale turbulence does not destroy the conditions of applicability of this approximation at intermediate scales, i.e. a separation of scales that we will discuss below.

The fact that the QG vorticity equation (Eq. 13) can be transformed into a conservation law of the pseudo potential vorticity (Eq. 14), similarly to the vorticity of 2-D turbulence, enabled Charney (1971) to argue that QG turbulence (QGT) should display, like 2-D turbulence, both a direct (downscale) enstrophy and an inverse (upscale) energy inertial range (Fjortoft, 1953; Kraichnan, 1967). Let us mention that both the expressions “inertial range” and “cascade” used by different authors (including in the present debate on the atmospheric dynamics at intermediate scales) denote the same physical phenomenon (nonlinear transfer of a given quantity across scales in a conservative manner), although with a slightly different emphasis on the possible underlying dynamical mechanisms.

More precisely, Charney claimed that there exists a mathematical isomorphism between the 2-D relative vorticity and the 3-D quasi-potential vorticity, although this is not straightforward as soon as the Brunt-Väisälä frequency is variable along the vertical (Herring, 1980) because whereas QG involves a 2-D advection operator, the advected quasi-potential vorticity is 3-D. Nevertheless, this theoretical possibility has been rather confirmed with the help of analytical closures (Herring, 2001, 1980; Salmon et al., 1978; Salmon, 1978) and numerical simulations (McWilliams et al., 1994; Hua and Haidvogel, 1986; Vallis, 1985; Fu and Flierl, 1978). However, Herring (2001) concluded that many questions remained open. In particular, he mentioned the empirical finding by Lindborg (1999) that the the spectrum $E(k) \approx k^{-5/3}$ thought to be a 2-D inverse energy inertial range by Gage (1979) and by Lilly and Paterson (1983) was rather a direct energy inertial range.

Some numerical evidence in the direction of a composite horizontal spectrum of the type $A k^{-5/3} + B k^{-5}$ was obtained by Tung and Orlando (2003), which is the unique reference issued by LTNCG to numerical models. However, this was obtained with only the help of a two-layer QG model (Welch and Tung, 1998), contrary to the simulations by Hua and Haidvogel (1986) that involved a large number of layers. With only two layers, not much can be done about either the boundary layer, therefore the boundary conditions, nor the horizontal/vertical anisotropy. The upper rigid boundary condition considered requires the introduction of Ekman damping whose exact choice would influence the large scale behaviour. Furthermore, Smith (2003) convincingly pointed out that the estimated “meso-scale” spectrum slope $\approx 5/3$ may well be spurious due to an artificial build up of enstrophy (and therefore of energy) at the smallest explicit model scales.

Overall, models do not seem to overcome the intrinsic limitations of the QG approximation. The first reason is that the QG approximation is fundamentally inappropriate for the mesoscale range: all the necessary approximations are no longer justified (e.g. the Rossby number becomes much larger than unity). Therefore, it seems unreasonable to hope that some 3-D-like behaviour would occur over the mesoscale range in a QG model, independently of the fact that it presumably occurs in nature. Secondly, the directions of the transfer in the inertial ranges can be theoretically inferred from inviscid statistical equilibria of the systems (Kraichnan, 1971), which yield an inverse energy and a direct enstrophy inertial range for both 2-D and QG turbulence. Therefore, we can safely conclude – contrary to LTNCG – that there is neither theoretical argument nor model evidence in favour of two direct inertial ranges. Let us mention that the
fact that pressure coordinates are commonly used in meteorology – as advocated by LTNCG – does not prevent them from theoretically introducing biases in statistical analyses because their possible dynamical significance strongly depends on the validity of a number of approximations that we have brought into question above.

It is worthwhile noting that a QG model cannot give insights on the separation of scales that would ensure its own physical relevance. This separation of scales between a 2-D regime and a 3-D regime, if it existed, could be easily destabilised by the vortex stretching mechanism, and we have already discussed the fact that the QG approximation greatly modifies the corresponding vector in the vorticity equation. More generally, small scale 3-D turbulence cannot be understood as only dissipating large scale structures by eddy viscosity, because it also generates larger scale structures by backscattering (Lesieur and Schertzer, 1978) or renormalized forcing (Forster et al., 1977; Fournier and Frisch, 1983). The zigzag instability (Billant, 2010; Billant et al., 2010), which bends 2-D columnar vortices and introduces a full 3-D regime is rather illustrative of the instability of 2-D turbulence to 3-D perturbations. The crucial importance of the separation of scales for numerical weather forecasts was explicitly stated and discussed by Monin (1972) and it explains why the concept of “mesoscale gap” (Van der Hoven, 1957) was so cherished during the early history of weather forecasting, this despite its existence having been criticised on observational grounds by Robinson (1967). This question of scale separation must not be reduced to a problem of boundary layer modelling as suggested by LTNCG, because this problem must be solved over all horizontal levels. We already mentioned that Charney was well aware of this problem.

LTNCG are right to point out that in QG it is the sum of potential and kinetic energy that takes the place of the kinetic energy. However, this cannot solve the problem of vertical statistical exponents being different from the horizontal ones. Indeed, the naive scaling exponents, i.e. without including intermittency effects or logarithmic corrections (Kraichnan, 1967), obtained by arguments à la Kolmogorov (1941) are ultimately based on dimensional analysis of the relevant fluxes. Therefore in order to get different horizontal and vertical scaling exponents (scaling anisotropy) one needs to consider a vertical turbulent flux with physical dimensions different from energy. The theoretical choice of the buoyancy force variance flux (Schertzer and Lovejoy, 1984, 1985b) still seems reasonable, although it was a bit bold when first proposed. Indeed, since the pioneering studies of Adelfang (1971) and Endlich et al. (1969) and especially since the 1980s, there has been a growing body of evidence that the scaling is indeed anisotropic with at least roughly the predicted exponents, see the numerous references cited in our previous comment (Lovejoy et al., 2009b).

4 Direct scaling analysis and Generalized Scale Invariance

To avoid the limitations induced by the QG approximation, we must proceed to a direct scaling analysis of the Navier-Stokes equations. We first recall how it can be done, when there is no rotation and/or gravitational field, in a formal manner which is both more straightforward and more general than a spectral analysis. This corresponds to analysing how the equations could be invariant under a space contraction/dilation, i.e. how the different fields (velocity, pressure, etc.) can balance at all scales. This is therefore quite distinct from a scale analysis, where only a few relevant fields at a given scale are preserved. This is obtained by analysing the effect of a space contraction/dilation on each field, something that is classically done for isotropic space contraction/dilation, whereas stratified and rotating flows require a generalisation to strongly anisotropic space contraction/dilations, e.g. they scale differently according to space directions, as illustrated by Fig. 1. These contraction/dilations were considered in the framework of Generalised Scale Invariance (GSI, Schertzer and Lovejoy, 1985b) that was first presented at the AMS Charney Memorial conference (Boston, 1983). In this section, we introduce an original technique to systematically study the relevance of such transforms in nonlinear equations.

In the present case, we want to reconcile the facts that vorticity at large scales is dominated by Earth’s rotation and that the nonlinearity of the vorticity equation generates scaling by transferring this large scale input across a wide range of scales. We are therefore led to a generalisation of the classical notion of spontaneous symmetry breaking. The latter corresponds to the fact that solutions of a system having a given symmetry, e.g. rotational symmetry for the vorticity equations, do not in general respect individually this symmetry, whereas they usually do it statistically. A classical example is the buckling of a cylindrical pillar under an increasing axial load: the pillar finally bends towards some definite direction rather than to being cylindrically deformed. With the help of such anisotropic space transforms, we generate a statistical break of the statistical rotational symmetry of the vorticity equations to obtain a new set of equations whose symmetry corresponds to anisotropic scaling. As discussed below, to systematically study the effect of space contraction/dilations, it is convenient to use the general notion of a “pullback” transform of a field by these contraction/dilations.

When there is neither rotation nor gravitational field, the Navier-Stokes equations (Eq. 1) and the mass conservation equation (Eq. 2) are rotationally invariant, as well as formally invariant under isotropic contraction/dilation $T_\lambda$ of space for any arbitrary scale ratio $\lambda$ ($\lambda > 1$ for a contraction, $\lambda < 1$ for a dilation):

$$x \mapsto T_\lambda x = x/\lambda$$  \hspace{1cm} (18)
It suffices indeed to suitably renormalize/rescale the other variables by various powers of this scale ratio \( \lambda \):

\[
t \mapsto t/\lambda^{1+\gamma}; u \mapsto u\lambda^\gamma; f \mapsto f\lambda^{2\gamma+1}; v \mapsto v\lambda^{\gamma-1}; \rho \mapsto \rho\lambda^\gamma
\]

(19)
defined by the singularities \( \gamma, \gamma' \) of respectively \( u, \rho \), see (Schertzer and Lovejoy, 2004) for further discussion in particular on the question of the dissipation and forcing terms.

For a unique singularity \( \gamma \), the corresponding power spectrum is a power-law \( E(k) \approx k^{-\beta} \) with the spectral slope \( \beta = 1 - 2\gamma \), because it scales like \( u^2 \). One thus obtains the celebrated Kolmogorov scaling of the velocity field for the (Kolmogorov’s) singularity \( \gamma_K = -1/3 \), which is defined by considering that the energy flux density \( \epsilon \), which scales like \( u^3/\lambda \), is homogeneous and strictly scale invariant in the inertial range (Kolmogorov, 1941).

We are now looking for a generalisation of this scaling analysis for anisotropic equations with the help of generalised contraction/dilation operators \( T_\lambda \), which still form a one-parameter multiplicative group \( (T_\lambda \circ T_\mu = T_{\lambda \mu}, \lambda, \mu \in \mathbb{R}^+ \) ), as in the isotropic case, and have therefore a generator \( G \). In the framework of linear GSI, which is also known under the name of “operator scaling”, \( G \) and \( T_\lambda \) are matrices:

\[
T_\lambda = \exp(-G \log \lambda) = \lambda^{-G}
\]

(20)

Obviously, \( G = \text{Id} \) (\( \text{Id} \) being the identity matrix) corresponds to the classical, scalar scaling (Eq. 18). It is now convenient to use the notion of a “pullback” transform (or “composition operator”, Shapiro, 1993) of a field \( u \) by a given space transform, here a contraction/dilation \( T_\lambda \). It is so general that it is often passed over without mention and it can be just seen as a convenient and compact notation for the scaling of a given field. As illustrated by Fig. 2, the pullback \( T_\lambda^* \) corresponds to a straightforward generalisation to (infinite dimensional) functional spaces of the (contravariant) change of coordinates on (finite dimensional) vector spaces:

\[
\forall x : T_\lambda^*(u)(x) = u(T_\lambda x)
\]

(21)

The composed function \( u(T_\lambda) \) pulls back the field \( u \) from the coordinates \( y \) to \( x \), with the “change of coordinates” \( y = T_\lambda(x) \) (for further discussion see Schertzer et al., 2010; Schertzer and Lovejoy, 2011). This transform can be extended for differential operators \( D \):

\[
\forall f : T_\lambda^*(D)T_\lambda^*(f) = T_{\lambda^*}^*(Df)
\]

(22)

For instance, the pullback of the gradient operator \( \nabla \) is:

\[
T_{\lambda^*}^*(\nabla) = T_\lambda^{-1}\nabla
\]

(23)
due to the fact that the transform \( T_\lambda \) is linear and is therefore its own Jacobian matrix. To study the (possible) anisotropy of atmospheric dynamics, it is sufficient to consider a diagonal generator under the following form:

\[
G = \text{diag}(g_i); g_1 = g_2 = 1; g_3 = H_\varepsilon = 1-h
\]

(24)

with an exponent \( 0 \leq H_\varepsilon \leq 1 \) that defines the anisotropy of \( T_\lambda \). For instance, \( H_\varepsilon = 0 \) corresponds to a (scaling) 2-D flow, \( H_\varepsilon = 1 \) to a 3-D (isotropic) scaling flow, and more generally intermediate values to a \( (2 + H_\varepsilon) \)-dimensional (anisotropic) scaling flow. Now, a generalised scaling of the field \( u(x,t) \) is defined by the fact that its pullback transform \( T_{\lambda^*}^* \) corresponds to (possibly random) generalised contraction with a generator \( \Gamma \):

\[
T_{\lambda^*}^* u = \lambda^\Gamma u
\]

(25)

The simplest case corresponds to the fact that both generators \( G \) and \( \Gamma \) commute, i.e. are diagonal on the same vector base. We consider the following particular case:

\[
\Gamma = \text{diag}(\gamma_i); \gamma_1 = \gamma_2 = \gamma; \gamma_3 = \gamma + h
\]

(26)

where \( \gamma \) is a given (scalar) singularity.

To give an illustration, let us consider the scaling model of a stratified atmosphere based on the conservation (i.e. strict scale invariance) of the energy flux along the horizontal and of the buoyancy force variance flux along the vertical (Schertzer and Lovejoy, 1984, 1985b). These conservation laws single out respectively the Kolmogorov’s singularity \( \gamma_K = -1/3 \) and the Bolgiano-Obukhov singularity \( \gamma_{BO} = -3/5 \). The latter is obtained similarly to the former, but for a flux that scales like \( u^3/\lambda^5 \) (Bolgiano, 1959; Obukhov, 1962). Note that these scaling exponents were both obtained assuming homogeneous and isotropic fluxes, whereas Schertzer and Lovejoy considered heterogeneous and anisotropic fluxes. In this anisotropic framework the theoretical value \( H_\varepsilon = 5/9 \) of the exponent of anisotropy merely results from the correspondence between the singularities of Kolmogorov and Bolgiano-Obukhov: \( \gamma_K = H_\varepsilon \cdot \gamma_{BO} \).

The choice of having the same \( h \) in Eqs. (24), (26) not only makes, regardless of the \( h \) value, the incompressibility condition scale invariant, but also ensures that the advection term and the material derivative have the same scalar scaling:

\[
T_{\lambda^*}^*(u \cdot \nabla) = \lambda^{\gamma+1} u \cdot \nabla; T_{\lambda^*}^*(D/Dt) = \lambda^{\gamma+1} D/Dt
\]

(27)
Because we consider an asymmetry between the horizontal and the vertical, it is useful both for the derivation and the expression of the fractional vorticity equations to decompose the fields and operators into horizontal and vertical components (with respective indices h and v), e.g. for the velocity field $u$ and gradient operator $\nabla$:

$$u = u_h + u_v; \nabla = \nabla_h + \nabla_v$$ (28)

rather than with respect to a three-dimensional basis $(e_1, e_2, e_3)$, e.g.:

$$u = (u^1, u^2, u^3); \nabla = (\partial_1, \partial_2, \partial_3);$$ (29)

One readily obtains the following scaling for the velocity field $u$ and gradient operator $\nabla$:

$$T^s_\lambda(u) = \lambda^\gamma (u_h + \lambda^h u_v); \quad T^s_\lambda(\nabla) = \lambda(\nabla_h + \lambda^{-h} \nabla_v)$$ (30)

Due to their difference of scaling (as confirmed below), we have to distinguish the horizontal vorticity components respectively yielded from vertical motions $(\sigma)$ and from horizontal motions $(\tau)$, see Fig. 3 for illustration:

$$\omega = \omega_v + \omega_h; \quad \omega_h = \sigma + \tau; \quad \omega_v \equiv \nabla_h \times u_h; \quad \sigma \equiv \nabla_v \times u_h; \quad \tau \equiv \nabla_v \times u_h$$ (31)

It is straightforward to obtain the respective scaling of the vorticity components from Eqs.(30–31):

$$T^s_\lambda(\omega_v) = \lambda^\gamma \omega_v; \quad T^s_\lambda(\sigma) = \lambda^{1+\gamma+h} \sigma; \quad T^s_\lambda(\tau) = \lambda^{1+\gamma-h} \tau$$ (32)

On the other hand, the stretching vector has the following decomposition:

$$s = ((\sigma + \tau) \cdot \nabla_h + \omega_v \cdot \nabla_v) (u_h + u_v)$$ (33)

that yields the following scaling $((\sigma \cdot \nabla_h) u_v = 0$ due to the asymmetry of $\sigma$):

$$T^s_\lambda(s) = \lambda^{2(1+\gamma)} [\lambda^h (\sigma \cdot \nabla_h) u_h + \lambda^{-h} (\tau \cdot \nabla_h + \omega_v \cdot \nabla_v) (u_h + \lambda^h u_v)]$$ (34)

Due to the fact that the vorticity equations should be satisfied for any $\lambda$, which appears with three distinct exponents $(2(1+\gamma)) (h, 0, -h)$, we obtain from Eqs.(27), (34) (ignoring for the present time the baroclinic vector) the following set of (barotropic) fractional vorticity equations:

$$D\sigma / Dt = (\sigma \cdot \nabla_h) u_h$$ (35)

$$D\tau / Dt = (\tau \cdot \nabla_h + \omega_v \cdot \nabla_v) u_h$$ (36)

$$D\omega_v / Dt = (\tau \cdot \nabla_h + \omega_v \cdot \nabla_v) u_v$$ (37)

which can be rewritten under the following coordinate form:

$$D\partial_2 u^3 / Dt = (\partial_2 u^3 \partial_1 - \partial_1 u^3 \partial_2) u^1 / Dt$$

$$D\partial_3 u^2 / Dt = (\partial_3 u^2 \partial_1 - \partial_1 u^2 \partial_3) u^1 / Dt$$

$$D(\partial_1 u^2 - \partial_2 u^1) / Dt = (\partial_1 u^2 \partial_3 - \partial_3 u^2 \partial_1 + \partial_3 u^1 \partial_2) u^3$$

These equations (either under the vector or coordinate form) correspond to splitting the dynamical equation of the horizontal vorticity $\omega_v$ into two separate equations, respectively for $\sigma$ (the horizontal vorticity produced by vertical motions) and $\tau$ (the horizontal vorticity produced by horizontal motions): this corresponds to the statistical break discussed above of the rotational symmetry of the original 3-D vorticity equations. Reciprocally, the latter equations are obtained by coupling again these equations, either under vector form (with $(\sigma \cdot \nabla_h) u_v = 0$ as mentioned earlier):

$$D(\sigma + \tau) / Dt = (\sigma + \tau) \cdot \nabla_h u_h + \omega_v \cdot \nabla_v u_h$$ (41)

$$D\omega_v / Dt = (\tau \cdot \nabla_h + \omega_v \cdot \nabla_v) u_v$$ (42)

or under the coordinate form:

$$D(\partial_2 u^3 - \partial_3 u^2) / Dt = \partial_2 u^3 \partial_1 u^1 - \partial_1 u^3 \partial_2 u^1 + \partial_1 u^2 \partial_3 u^1 - \partial_3 u^2 \partial_1 u^1$$ (43)

$$D(\partial_3 u^2 - \partial_1 u^3) / Dt = \partial_3 u^2 \partial_1 u^2 - \partial_1 u^2 \partial_3 u^2 + \partial_2 u^3 \partial_1 u^2 - \partial_1 u^3 \partial_2 u^2$$ (44)

$$D(\partial_1 u^2 - \partial_2 u^1) / Dt = \partial_1 u^2 \partial_3 u^3 - \partial_3 u^2 \partial_1 u^1 + \partial_3 u^1 \partial_2 u^3 - \partial_2 u^1 \partial_3 u^3$$ (45)

This confirms the fact that solutions of Eqs. (35–37) form the subset of solutions of Eqs. (43–45) that statistically respect the anisotropic scaling prescribed by Eqs. (24), (26) and statistically break therefore the isotropic scaling of Eqs. (43–45).
Whereas the QG linearisation of the stretching vector keeps only the evolution of the vertical vorticity (Eq. 37), with a unique source term \((\omega_v \cdot \nabla_v)u_v\) and with the further approximation \(\omega_v \simeq 2\Omega\), the fractional vorticity equations preserve all the nonlinearities. For instance, the term \((\omega_v \cdot \nabla_v)u_h\) (in Eq. 36) tilts part of the vertical vorticity \(\omega_v\) into the horizontal component \(\tau\), i.e. the latter does not necessarily remain zero if it was thus initially. The absence of a corresponding tilting term in the evolution equation of \(\sigma\) (Eq. 35) shows that \(\sigma\) and \(\tau\) have no symmetrical roles and it also emphasises the importance of the splitting of the dynamical equation of horizontal vorticity into two separate equations. However, both \(\sigma\) and \(\tau\) can undergo nonlinear stretching by (non-zero) horizontal gradients of the horizontal velocity (Eqs. 35–36), this implies in particular that solutions with \(\sigma = 0\) are unstable. These features are illustrative of the huge difference between scale analysis and scaling analysis: in the former case (QG) all the source terms but one were cancelled in the vorticity equations, whereas in the present anisotropic scaling analysis, they are all preserved but no longer contribute in an isotropic manner. For instance, the interactions of \(\sigma\) and \(\tau\) with the horizontal gradient of the horizontal velocity become differentiated (with separate equations Eqs. 35–36), whereas only their sum intervenes in the 3-D vorticity equations (Eq. 41).

Although our paper is focused on the barotropic case, let us quickly show that the baroclinicity can be readily taken into account at all scales. Indeed, the baroclinic vector \(b\) has the following pullback:

\[
T^*_k(b) = \lambda^{2(v+1)}(\lambda^{-h} b_h + b_v) \tag{46}
\]

One may note it has an opposite behaviour to that of the potential vector \(\psi\):

\[
T^*_k(\psi) = \lambda^{\sigma-1}(\lambda^h \psi_h + \psi_v) \tag{47}
\]

which corresponds to the expected phenomenology: to be aligned along the vertical for large scales, whereas it is no longer the case for smaller scales. The comparison of the scaling of the vorticity components (Eq. 32) and of the baroclinic vector (Eq. 46) shows that baroclinicity does not affect the horizontal vorticity generated by vertical motions of \(\sigma\), a fact that emphasises its particularity with respect to \(\tau\) and \(\omega_v\). The barotropic fractional vorticity equations (Eqs. 35–37) yield therefore the corresponding baroclinic equations:

\[
D\sigma / Dt = (\sigma \cdot \nabla_h)u_h \tag{48}
\]

\[
D\tau / Dt = (\tau \cdot \nabla_h + \omega_v \cdot \nabla_v)u_h + b_h \tag{49}
\]

\[
D\omega_v / Dt = (\tau \cdot \nabla_v + \omega_v \cdot \nabla_v)u_v + b_v \tag{50}
\]

The fractional vorticity equations (either barotropic, Eqs. 35–37) or baroclinic (Eqs. 48–50) were derived from the usual vorticity equations with the assumption of a given anisotropic scaling behaviour (Eqs. 24–26) of particular solutions. An inspection of these derived equations shows that conversely their solutions should statistically respect this type of anisotropy, however there is no hint that the theoretical value \(H_\tau = 5/9\) plays a special role. Indeed, this theoretical value was obtained in fact with the help of thermodynamic considerations, whereas we explained that we postpone to future works the treatment of the energy evolution equation in a similar manner to the way we treated the vorticity equations. On the other hand, it is worthwhile noting that the obtained fractional vorticity equations already correspond to an achievement that goes well beyond the question of the atmospheric dynamics scaling. Indeed, a precise definition of \(d\)-dimensional turbulence for non integer \(d\)'s has only been addressed in an isotropic and statistical framework (Fournier and Frisch, 1978) for the (average) energy spectrum and with the help of a second-order closure, whereas we obtained (apparently) deterministic, dynamical equations for the field itself of the vorticity, and therefore of the velocity. Furthermore, the fractional part \(H_\tau\) of this non integer dimension is physically meaningful: \(H_\tau\) measures the extent of the vertical motions, in fact the fractional depth of the corresponding pancake turbulence, from 0 (flat 2-D turbulence) to 1 (3-D turbulence).

5 Conclusions

We believe that the critical analyses of the scale analysis leading to the QG approximation (Sect. 2) and of the scaling analysis of the resulting QGT (Sect. 3) first rule out LT-NCG’s claim that the significance of the spectrum power law \(E(k) \approx k^{-3}\) on scales \(\geq 600\) km obtained from the GASP and Mozaic databases could not be brought into question because it corresponds to a QGT prediction. Beyond this necessary clarification, we furthermore put forward an alternative (Sect. 4) to the QG approximation that reconciles the facts that Earth’s rotation dominates vorticity at large scales and that the vorticity stretching nonlinearly transfers this large scale input over a wide range of scales in a scaling manner. Indeed, by introducing a statistical anisotropic break of the statistical isotropic scaling of the vorticity equations, we obtain fractional vorticity equations, whose solutions should statistically satisfy this anisotropic scaling over a wide range of scales and therefore define a turbulence of (fractal) dimension \(D = 2 + H_\tau\) \((0 \leq H_\tau \leq 1)\). This corresponds to the dynamical step towards this alternative to QG, which should be completed by a similar treatment of the thermodynamic energy evolution equation. Overall, our claim that atmospheric turbulence has a dimension \(D = 2 + H_\tau\), presumably with the theoretical value \(H_\tau = 5/9\), and is therefore so much distinct from quasi-2-D and quasi-3-D turbulences, seems to result from the application of “Occam’s Razor” in the framework of a fruitful dialogue between theories and data, as defined by Medawar (1969): “scientific reasoning is a kind of dialogue
between the possible and the actual, between what might be and what is in fact the case”.

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